UNIVERSITY OF CAMBRIDGE INTERNATIONAL EXAMINATIONS

GCE Advanced Subsidiary Level and GCE Advanced Level

MARK SCHEME for the October/November 2010 question paper for the guidance of teachers

9702 PHYSICS

9702/41

Paper 4 (A2 Structured Questions), maximum raw mark 100

This mark scheme is published as an aid to teachers and candidates, to indicate the requirements of the examination. It shows the basis on which Examiners were instructed to award marks. It does not indicate the details of the discussions that took place at an Examiners' meeting before marking began, which would have considered the acceptability of alternative answers.

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Section A

1	(a)	force per unit mass	(ratio idea essential)	B1 [1]
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(b) graph: correct curvature M1 from
$$(R, 1.0 g_s)$$
 & at least one other correct point A1 [2]

(c) (i) fields of Earth and Moon are in opposite directions

either resultant field found by subtraction of the field strength

or any other sensible comment

so there is a point where it is zero

(allow
$$F_E = -F_M$$
 for 2 marks)

(ii)
$$GM_E/x^2 = GM_M/(D-x)^2$$
 C1
 $(6.0 \times 10^{24})/(7.4 \times 10^{22}) = x^2/(60R_E-x)^2$ C1
 $x = 54R_E$ A1 [3]

(iii) graph:
$$g = 0$$
 at least $\frac{2}{3}$ distance to Moon
$$g_{\rm E} \text{ and } g_{\rm M} \text{ in opposite directions}$$

$$\text{correct curvature (by eye) and } g_{\rm E} > g_{\rm M} \text{ at surface}$$

$$\text{A1} \quad [3]$$

- 2 (a) (i) no forces (of attraction or repulsion) between atoms / molecules / particles B1 [1]
 - (ii) sum of kinetic and potential energy of atoms / molecules M1 due to random motion A1 [2]
 - (iii) (random) kinetic energy increases with temperature no potential energy (so increase in temperature increases internal energy)

 A1 [2]
 - (b) (i) zero A1 [1]

(ii) work done =
$$p\Delta V$$
 C1
= $4.0 \times 10^5 \times 6 \times 10^{-4}$
= 240 J (ignore any sign) A1 [2]

(iii)

change	work done / J	heating / J	increase in internal energy / J
$\begin{array}{c} P \rightarrow Q \\ Q \rightarrow R \\ R \rightarrow P \end{array}$	+240 0 -840	-600 +720 +480	-360 +720 -360

(correct signs essential)
(each horizontal line correct, 1 mark – max 3)

B3 [3]

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3	(a)	(i)	resoi	nance		B1	[1]
		(ii)	ampl	A1	[1]		
	(b)	(i)	a =	$(-)\omega^2 x$ and $\omega = 2\pi f$ $4\pi^2 \times 4.6^2 \times 16 \times 10^{-3}$ $13.4 \mathrm{m s^{-2}}$		C1 C1 A1	[3]
		(ii) $F = ma$ = $150 \times 10^{-3} \times 13.4$					
				2.0N		A1	[2]
				ys 'below' given line and never zero t 4.6 Hz (or slightly less) and flatter		M1 A1	[2]
4	(a)	cha	rge / ¡	potential (difference) (ratio must be clear)		B1	[1]
	(b)	(i)	V = (Q / $4\pi \varepsilon_0 r$		В1	[1]
	((ii)	C = 0 so C	$Q/V = 4\pi \varepsilon_0 r$ and $4\pi \varepsilon_0$ is constant $\propto r$		M1 A0	[1]
	(c)	(i)	r = (6	$6./4\pi\varepsilon_0 r$ $6.8 \times 10^{-12}) / (4\pi \times 8.85 \times 10^{-12})$ $\times 10^{-2}$ m		C1 C1 A1	[3]
	((ii)		$CV = 6.8 \times 10^{-12} \times 220$ = 1.5×10^{-9} C		A1	[1]
	(d)	(i)	V = 0 = 83	$Q/C = (1.5 \times 10^{-9}) / (18 \times 10^{-12})$		A1	[1]
	((ii)	eithe	er energy = $\frac{1}{2}CV^2$ $\Delta E = \frac{1}{2} \times 6.8 \times 10^{-12} \times 220^2 - \frac{1}{2} \times 18 \times 10^{-12} \times 83^2$		C1 C1	
			or	= $1.65 \times 10^{-7} - 6.2 \times 10^{-8}$ = 1.03×10^{-7} J energy = $\frac{1}{2}$ QV $\Delta E = \frac{1}{2} \times 1.5 \times 10^{-9} \times 220 - \frac{1}{2} \times 1.5 \times 10^{-9} \times 83$ = 1.03×10^{-7} J		A1 (C1) (C1) (A1)	[3]

Mark Scheme: Teachers' version

Syllabus

Paper

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5	(a)	field	into	(the plane of) the paper		B1	[1]
	(b)		/ r = = (20	e to magnetic field <u>provides</u> the centripetal force Bqv 0 × 1.66 × 10 ⁻²⁷ × 1.40 × 10 ⁵) / (1.6 × 10 ⁻¹⁹ × 6.4 × 10 ⁻¹⁹ × 6.4 × 10 ⁻¹⁹ × 6.4 × 10 ⁻¹⁹	²)	B1 C1 B1 A0	[3]
	(c)	(i)	<u>sem</u>	icircle with diameter greater than 12.8 cm		B1	[1]
		(ii)	new	flux density = $\frac{22}{20} \times 0.454$		C1	
				$B = 0.499 \mathrm{T}$		A1	[2]
6	(a)	(i)	e.g.	prevent flux losses / improve flux linkage		B1	[1]
		-	e.m.	in core is changing f. / current (induced) <u>in core</u> ced current in core causes heating		B1 B1 B1	[3]
	(b)			value of the direct current producing same (mean) pov resistor	ver / heating	M1 A1	[2]
		(ii)	•	er in primary = power in secondary $I_S = V_S I_S$		M1 A1	[2]
7	(a)	(i)	e.g.	electron / particle diffraction		B1	[1]
		(ii)	e.g.	photoelectric effect		B1	[1]
	(b)	(i)				A1	[1]
			$\lambda = I$	age in energy = 4.57×10^{-19} J ac / E $63 \times 10^{-34} \times 3.0 \times 10^{8}$) / (4.57×10^{-19})		C1	
			= 4.4	$4 \times 10^{-7} \mathrm{m}$		A1	[2]
8	(a)	-	_	of a heavy nucleus (not atom/nuclide) (lighter) nuclei of approximately same mass		M1 A1	[2]
	(b)	¹ ₀ n ⁴ ₂ He ⁷ ₃ Li	;	(allow 4_2lpha)		M2 A1	[3]
	(c)	rang lose	je of kine	particles have kinetic energy particles in the control rods is short / particles stopped tic energy in rods	in rods /	B1 B1 B1	
		kinetic energy of particles converted to thermal energy					[3]

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Section B

Р	age 6	Mark Scheme: Teachers' version		Syllabus	Paper	•
		GCE AS	S/A LEVEL – October/November 2010	9702	41	
	(ii) use an amplifier coupled to the microphone (repeater amplifiers scores no mark)				M1 A1	[2]
12 (a	satellite i signal ar at a diffe different e.g. of fre	receives gre nplified and rent (carrier frequencies equencies u	nitted from Earth to satellite eatly attenuated signal transmitted back to Earth) frequency prevent swamping of uplink signal sed (6/4 GHz, 14/11 GHz, 30/20 GHz) any two other for additional physics)	(1) (1) (1) (1)	B1 B1	[4]
(b	advantag	e.g.	because orbits are much lower whole Earth may be covered in several orbits / with network		M1 A1 (M1) (A1)	
	disadvar	naye. e.y.	either must be trackedor limited use in any one orbitmore satellites required for continuous or	peration	M1 A1	[4]